

Spectroscopically resolving  
the 'inner' atmospheres  
of elderly stars

= TIGvival =

Mira and her sisters



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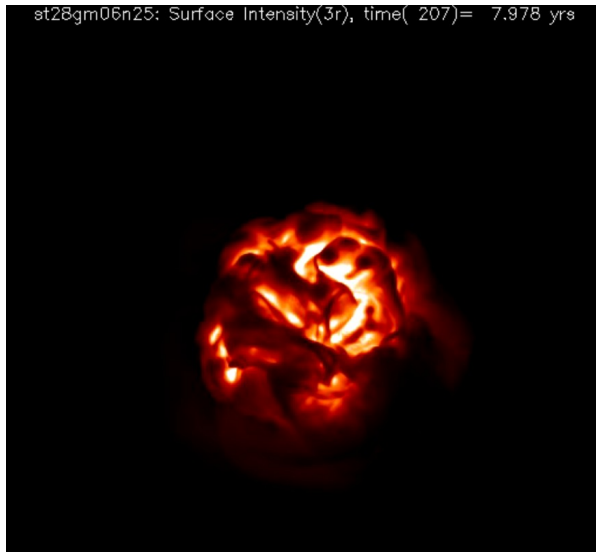
B. Aringer, B. Freytag, D. Engels, S. Bladh, S. Uttenthaler, R. Zamanov

- **AGB** stars - Long-Period Variables (**LPV**) - 'Miras'
- **TIGvival**: Mira and other AGB stars – 2013 until today
- **TIGRE/HEROS**: 'Heroes and villains'

Mass-loss mechanisms of

= AGB Stars =

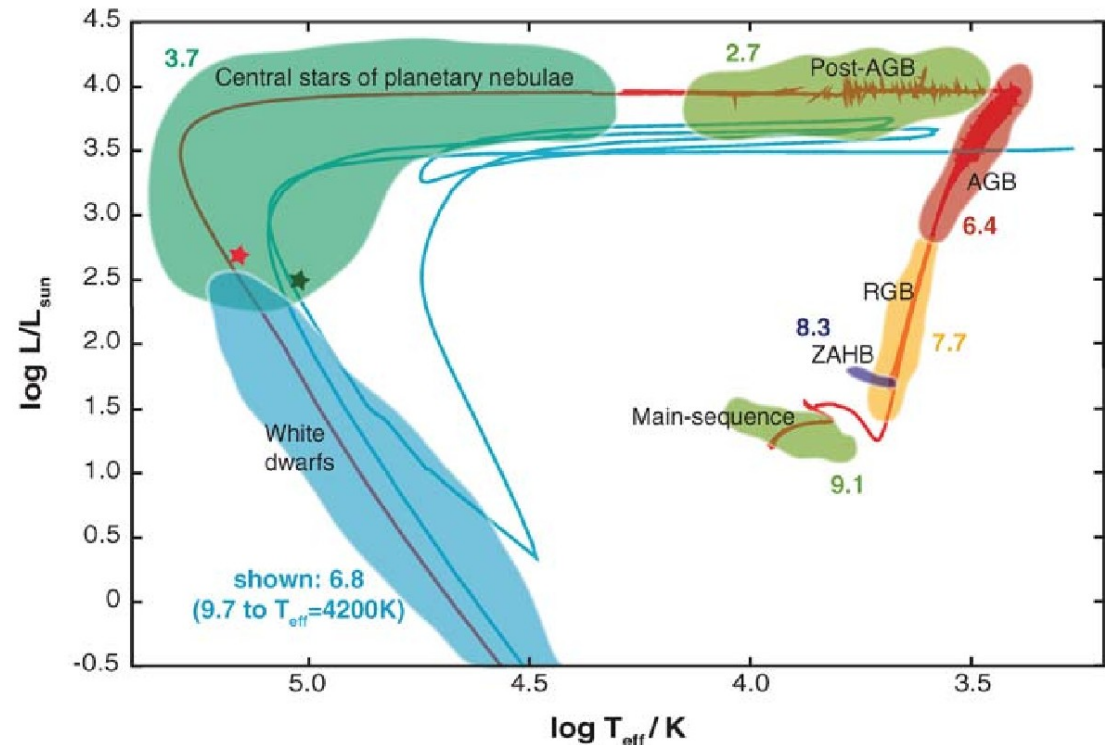
## 1. A one solar mass star loses on the order of 25% mass on the AGB



## 2. Snapshot of a 3D model: 1 solar mass with solar metallicity

(Freytag & Höfner 2017, A&A 600)

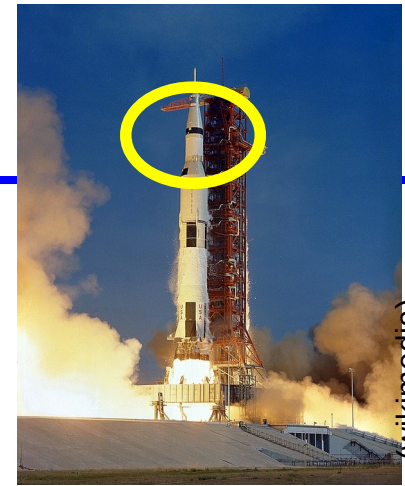
<http://www.astro.uu.se/~bf/movie/AGBmovie.html>



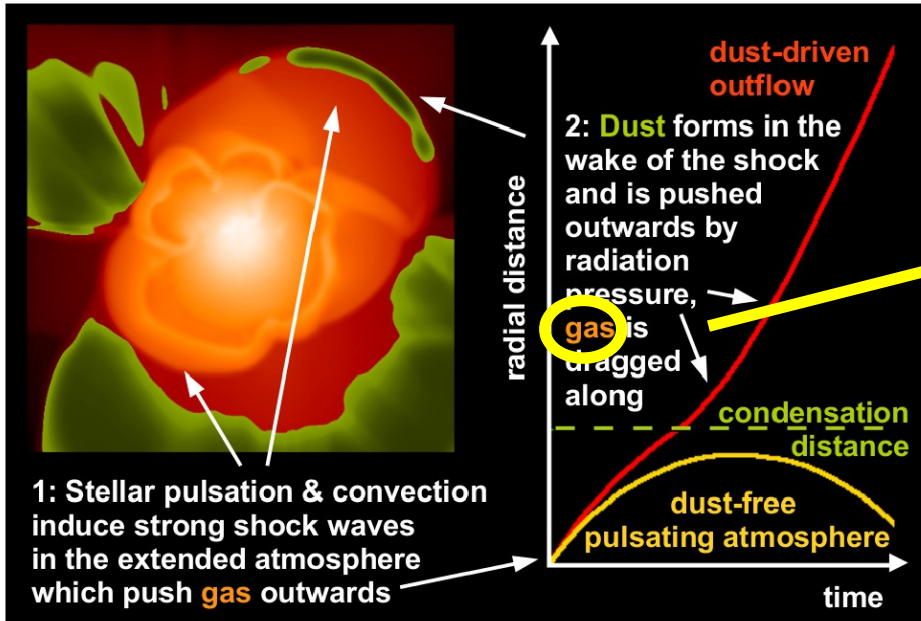
## 3. AGB initial mass ~0.5 to ~8 solar masses (2 solar masses with solar metallicity shown)

(Herwig 2005, ARA&A 43 – log t [yr] given)

# Mass loss of AGBs: PEDDRO



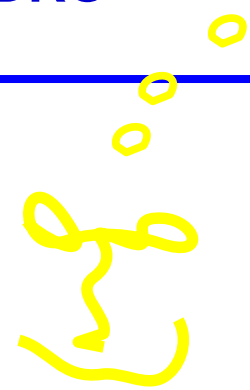
(wikimedia)



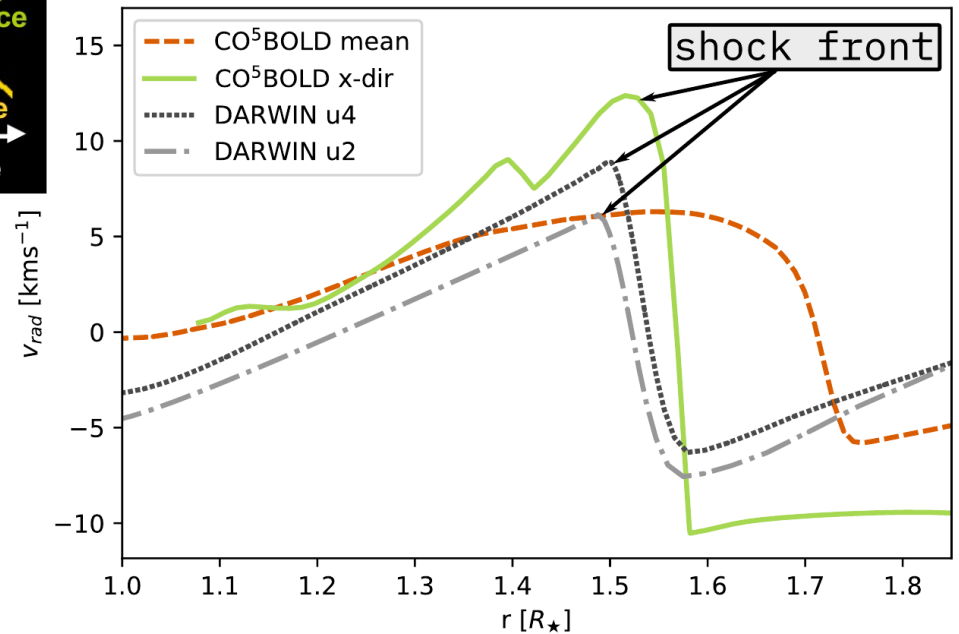
(Höfner 2016, CS 19)

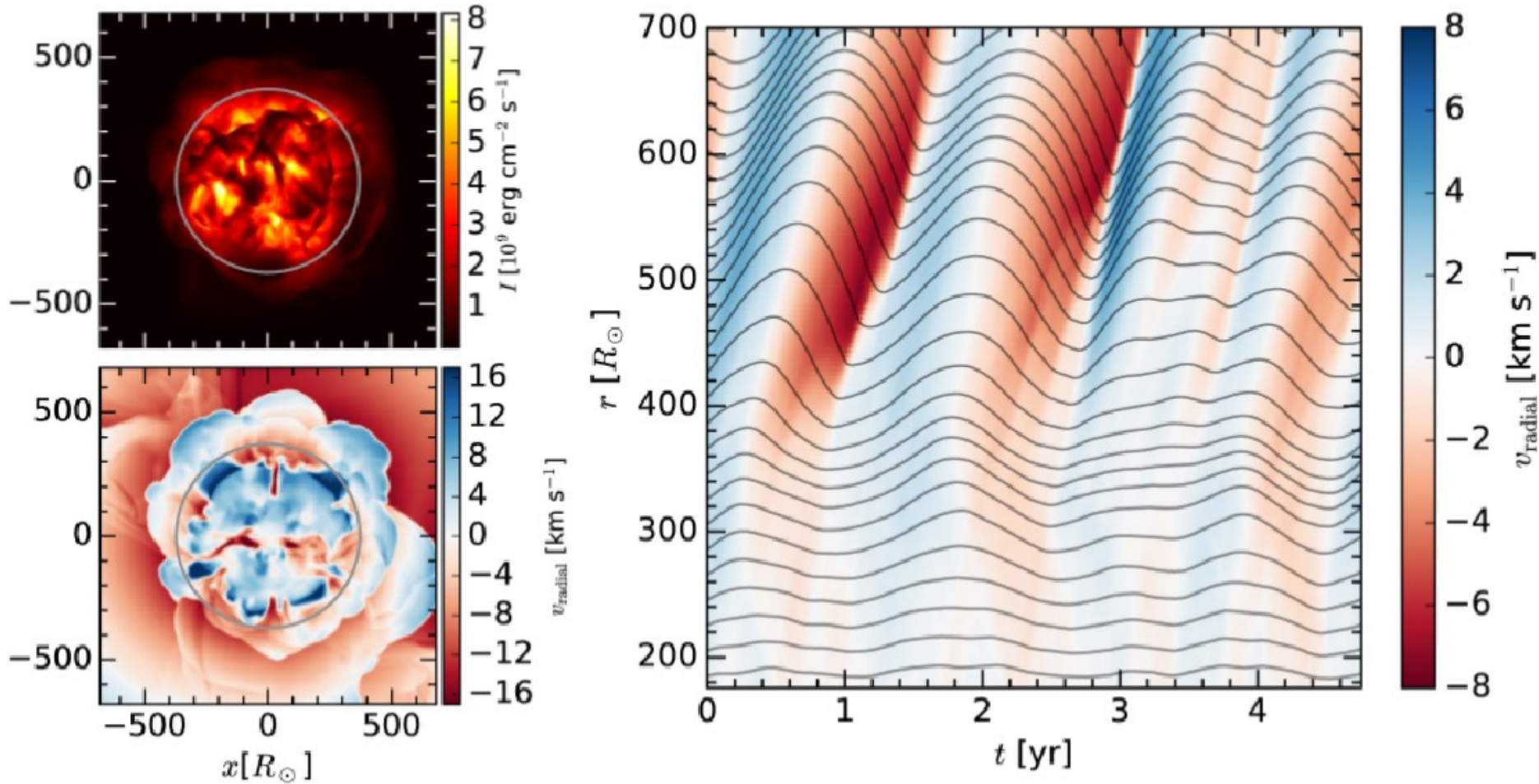
**PEDDRO = Pulsation-Enhanced Dust-Driven Outflow**

(Höfner & Olofsson 2018, AARv 26)



(Liljegren+ 2018, A&A 619)





(Höfner 2016, CS 19; Freytag & Höfner 2017, A&A 600; Liljegren+ 2018, A&A 619)

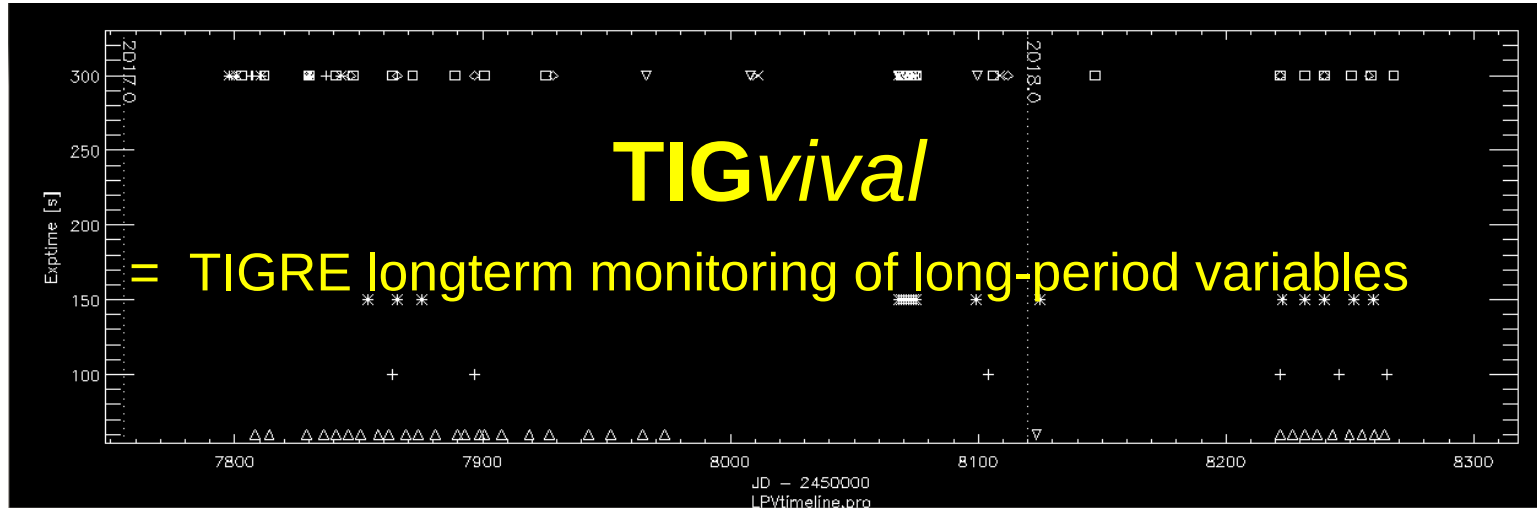




# = TIGvival =

(wikimedia)

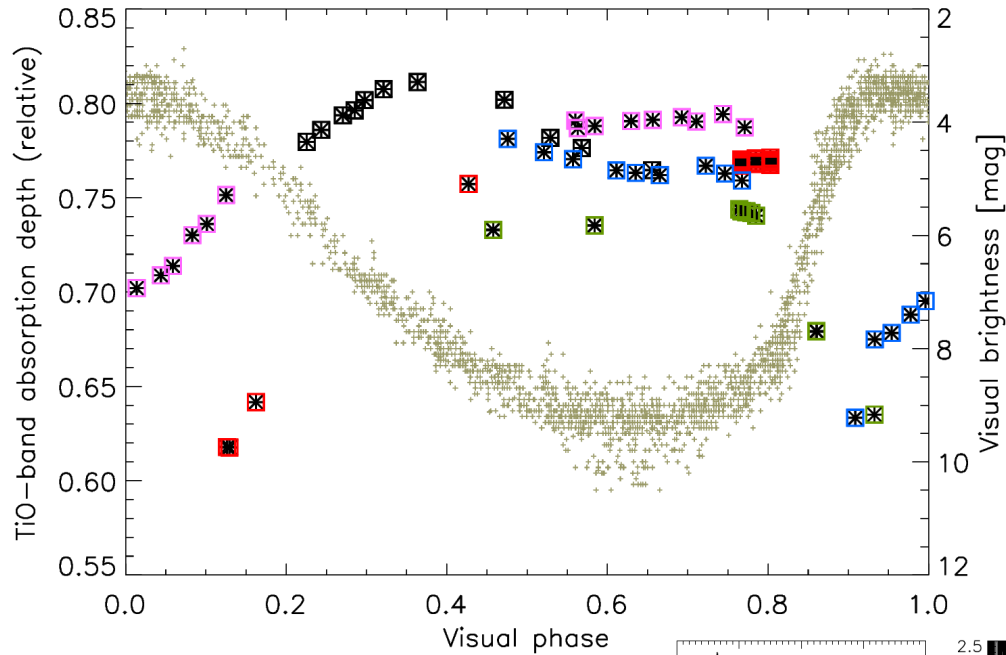
## Sleuthing Long-Period Variables (LPV)



Object	$m_V$
U Hya	5.4 – 4.6 (Per= 183 d.
G Her	5.5 – 4.3 (89 d, obs. 1
R Hya	8.5 – 5.5 (380 d, obs.
BK Vir	8.8 – 7.3 (150 , obs. 8
W Hya	10.0 – 5.8 (361 d, obs.
R Lep	10.0 – 5.5 (445 d, obs.
o Cet	10.2 – 2.0 (332 d, obs.
R Leo	11.0 – 4.5 (P= 310 d,
R Aql	11.0 – 5.5 (P= 271 d,
R Vir	11.5 – 6.0 (P= 146 d,

Object	Ra	Dec	$m_V$
U Ari	031103.0	+144800.3	8.0
S CMi	073243.1	+081905.2	8.0
X Hya	093530.3	-144128.6	8.5
R Leo	094733.5	+112543.8	5.5
R Tau	042818.0	+100944.8	9.0
U Ori	055549.2	+201030.7	10.0
R LMi	094534.3	+343042.8	7.5
X Oph	183821.1	+085002.8	7.0
ST Sgr	190129.2	-124534.1	9.0
S CrB	152124.0	+312202.6	7.5





## TIGvival: TiO 7055 Å in Mira

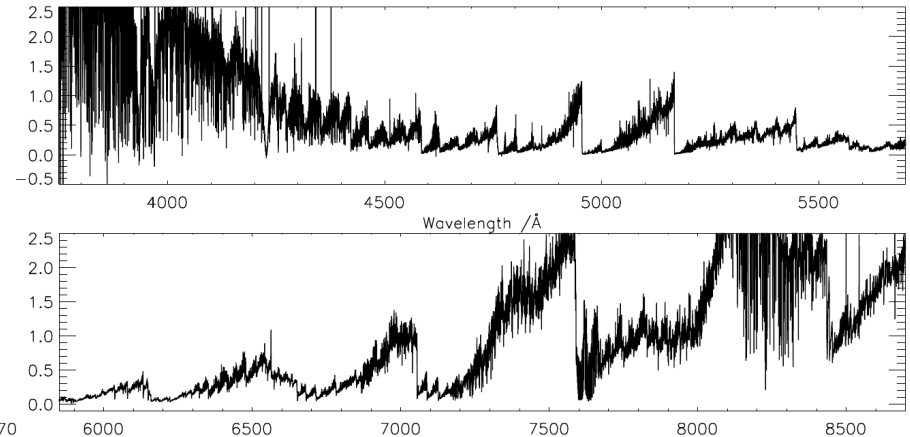
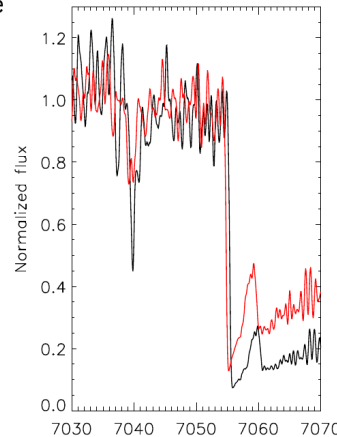
2013/14 (black) to 2018/19 (violet)  
+ AAVSO visual photometry

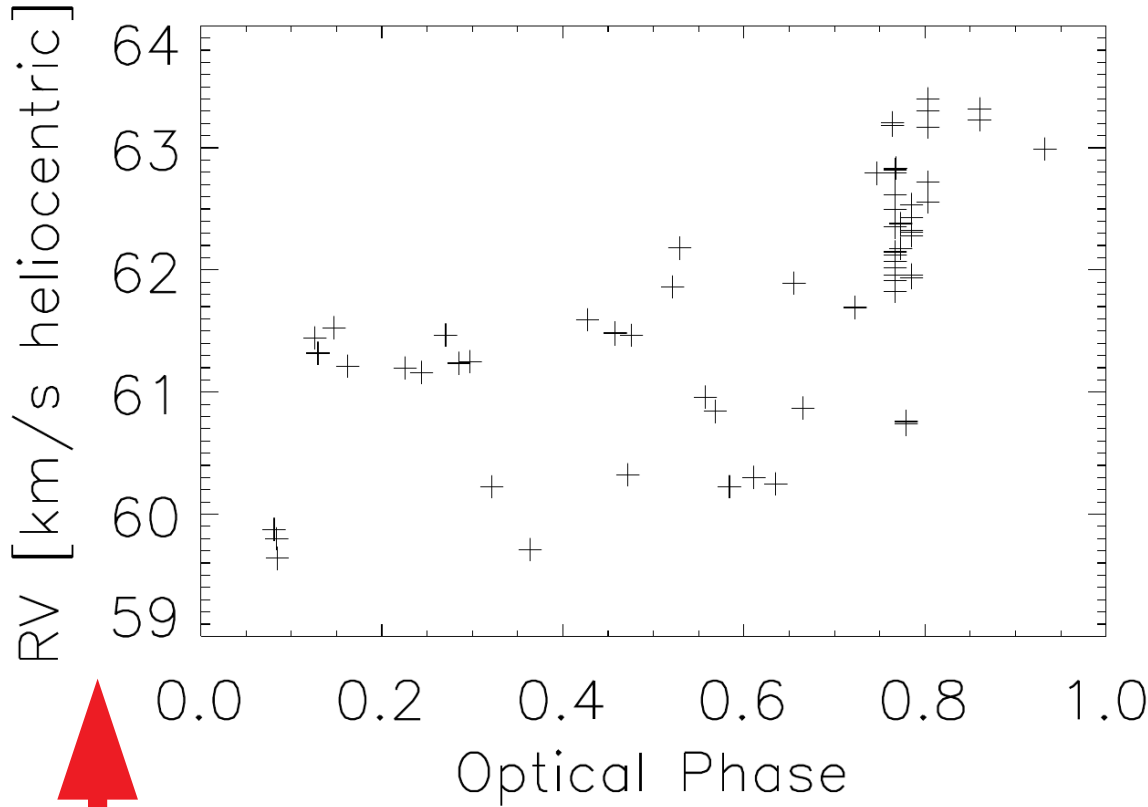
(cf. also Castellaz+ 2000, AJ 120)

## TIGRE spectra of Mira

$\Phi=0.16$  (red) and  $0.30$  (black)  
i.e. ~2 months apart

(Wolter+ 2019, IAUS 343,  
cf. also Bladh+ 2015, A&A 575)



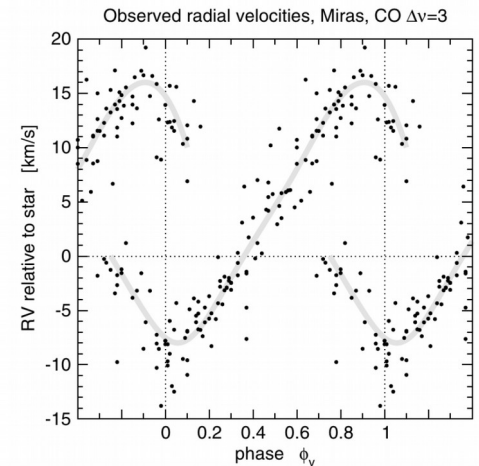


Mira's maser center velocity (H<sub>2</sub>O, CO, SiO)

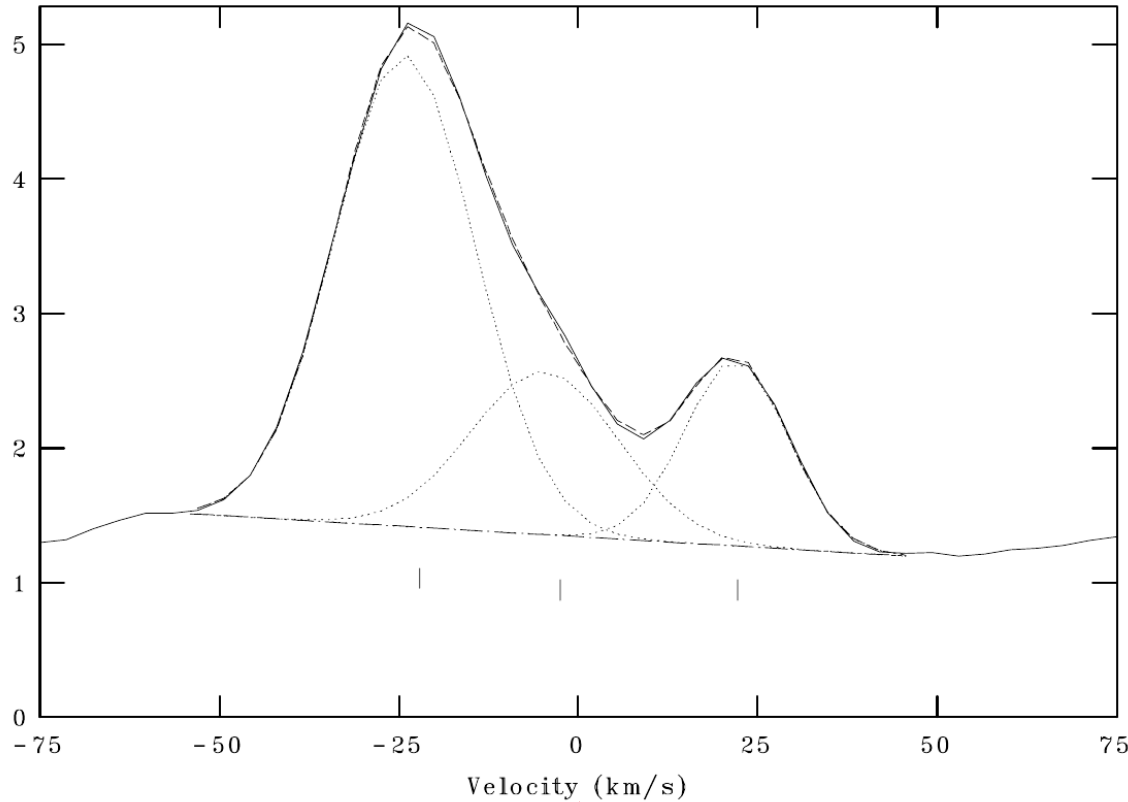
**~ 55 km/s** (heliocentric)

**TiO 7055 Å**  
Bandhead velocities in Mira  
(TIGRE)

**CO (17µm)**  
compiled for Mira variables

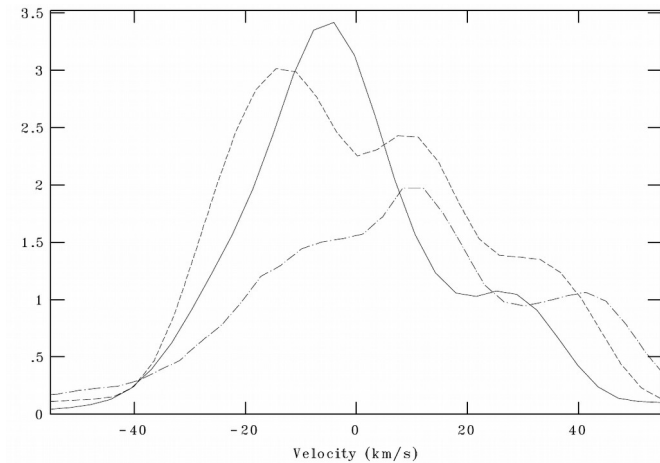


(Nowotny+ 2010, A&A 514)



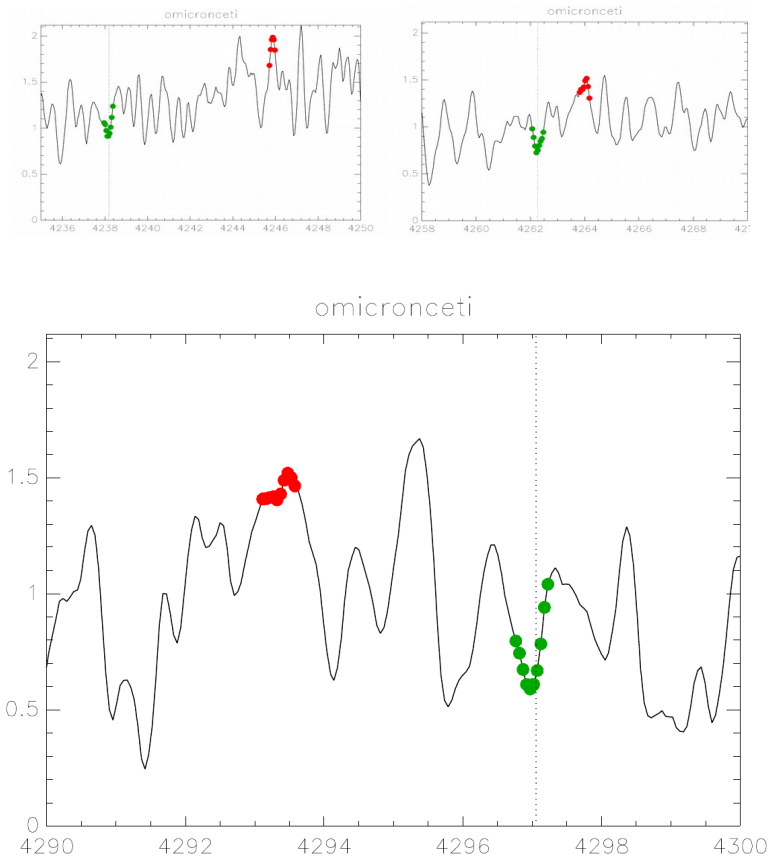
↑ ~ 55 km/s

H alpha velocities in Mira near phase 0 (TIGRE)

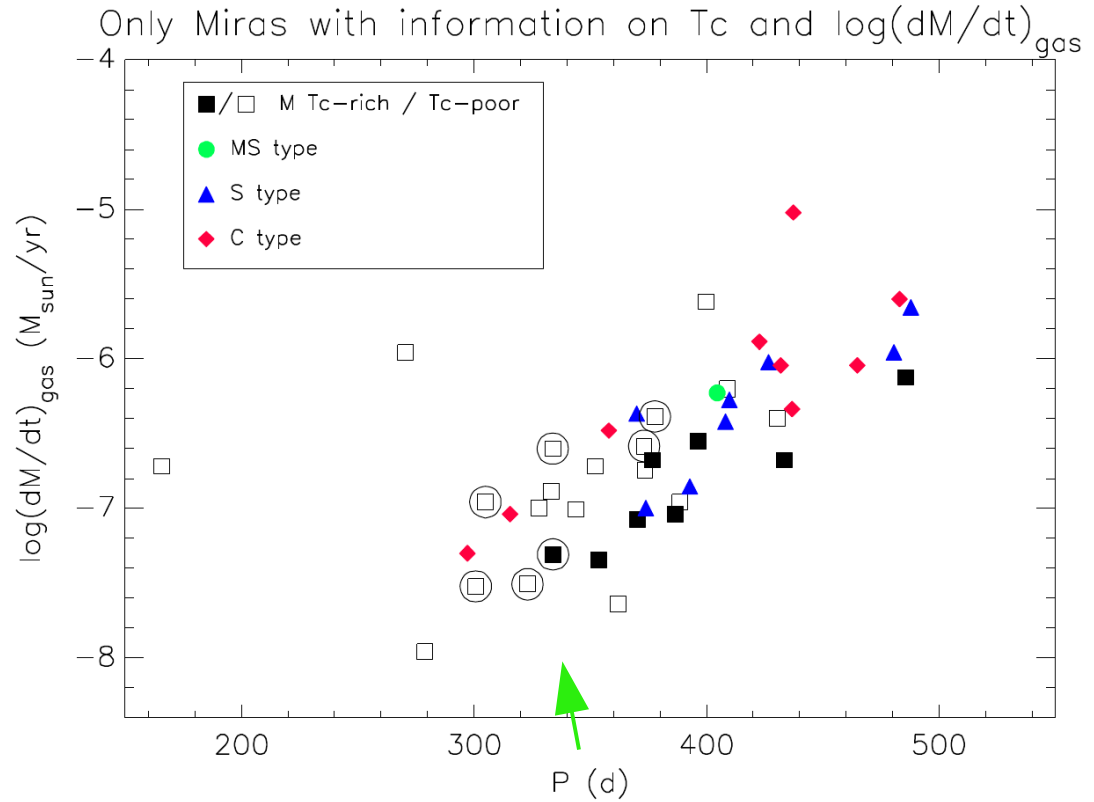


H delta, H gamma and H9 in Mira near phase 0.5

(cf. also Pierce, Willson, Beavers 1979, PASP 91)



**Tc lines in Mira**  
(TIGRE, S. Uttenthaler)



**Mass-loss and Tc in LPVs**  
(TIGRE measurements marked by O, S. Uttenthaler)

(cf. also Uttenthaler+ 2019, A&A 622)

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For >7 years, **TIGvival** will monitor ~10 long-period-variables to e.g. help gauge LPV-models

o **Cet (Mira)** @TIGRE exhibits quasi(?)-regular velocity and abundance variations – related to shocks in its photosphere and its ‘chromosphere’

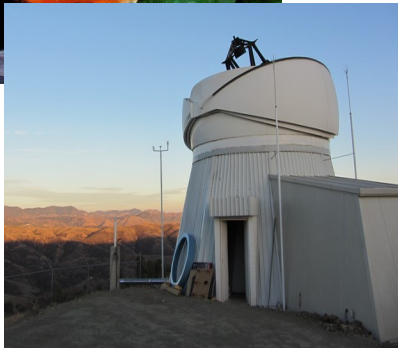
TIGRE/HEROS does measure at ~1 km/s precision and EW at a ~1% level – with careful calibration



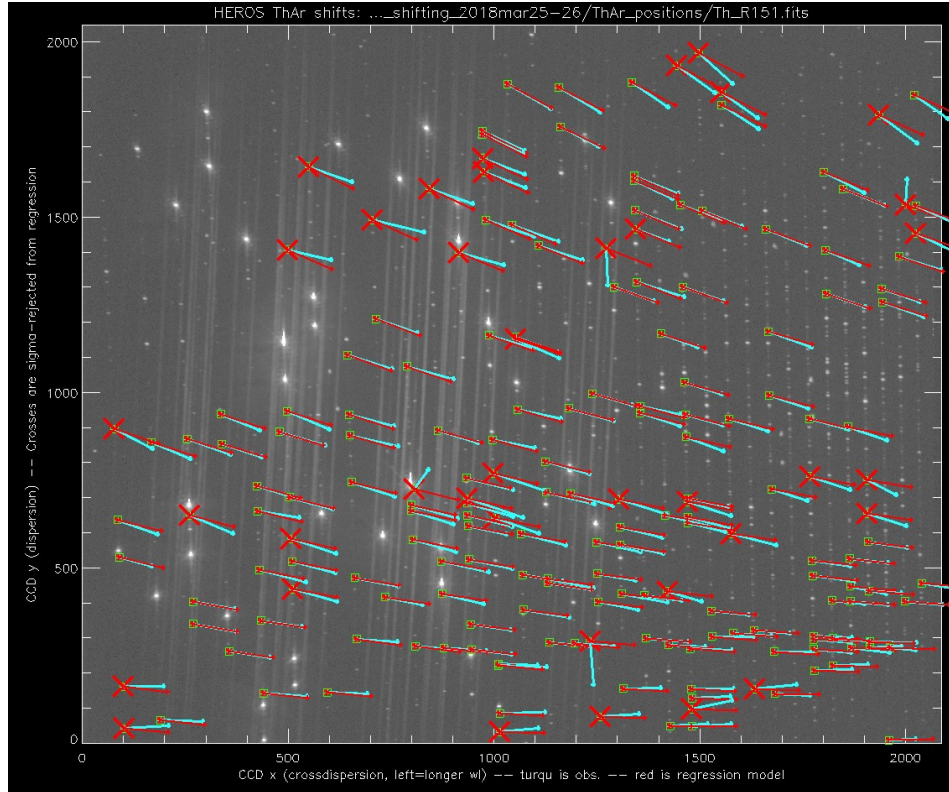
(wikimedia)



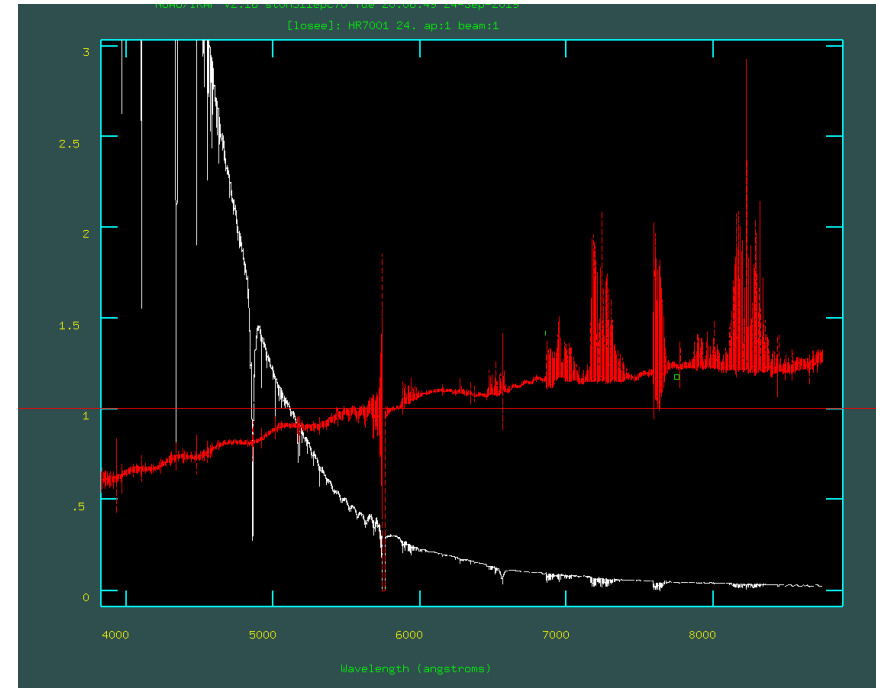
Painting: mira.com



(wikimedia)



**TIGRE/HEROS: ThAr at 15 vs. 20 °C  
(shifts x 100)**



**TIGRE/HEROS:  
Vega 1.2 vs. 3.0 arcsec**



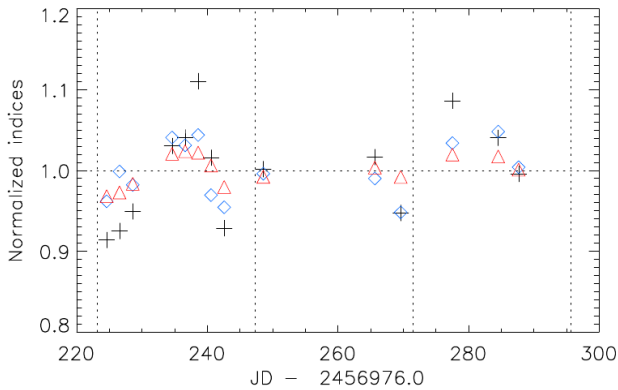


# APPENDIX



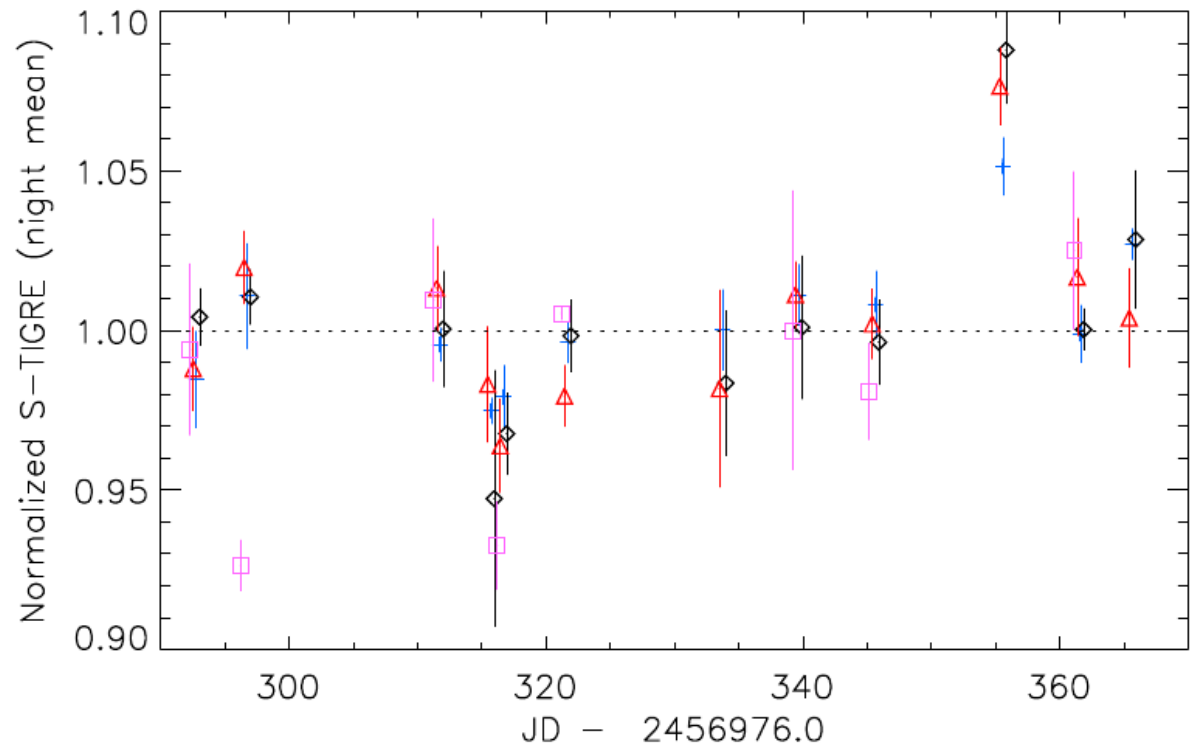


# OU And's chromospheric emission (2015 TIGRE + NARVAL)



**2015 NARVAL  
(June-August)**

(Borisova+ 2019, BlgAJ 31)



**2015 TIGRE (v3)  
(September-Nov.)**



EMPTY

